

#村上 豪¹⁾, 土屋 史紀²⁾, 山崎 敦¹⁾, 榎木谷 海¹⁾, 鍵谷 将人²⁾, 亀田 真吾³⁾, 吉岡 和夫⁴⁾, 桑原 正輝⁴⁾, 木村 智樹⁵⁾, 田所 彩華⁵⁾, 古賀 亮一⁶⁾, 清水 里香⁷⁾, 近藤 依央菜¹⁾, 伊庭 遼¹⁾

(¹⁾ 宇宙航空研究開発機構, (²⁾ 東北大学, (³⁾ 立教大学, (⁴⁾ 東京大学, (⁵⁾ 東京理科大学, (⁶⁾ 名古屋市立大学, (⁷⁾ 総合研究大学院大学

UV space telescope LAPYUTA: overview of the mission instruments and development updates

#Go Murakami¹⁾, Fuminori TSUCHIYA²⁾, Atsushi YAMAZAKI¹⁾, Umi Enokidani¹⁾, Masato KAGITANI²⁾, Shingo KAMEDA³⁾, Kazuo YOSHIOKA⁴⁾, Masaki Kuwabara⁴⁾, Tomoki KIMURA⁵⁾, Ayaka TADOKORO⁵⁾, Ryoichi KOGA⁶⁾, Riko SHIMIZU⁷⁾, Iona KONDOH¹⁾, Ryo IBA¹⁾

(¹⁾Japan Aerospace Exploration Agency, (²⁾Tohoku University, (³⁾Rikkyo University, (⁴⁾The University of Tokyo, (⁵⁾Tokyo University of Science, (⁶⁾Nagoya City University, (⁷⁾SOKENDAI

The Life-environmentology, Astronomy, and Planetary Ultraviolet Telescope Assembly (LAPYUTA) mission aims to carry out spectroscopy with a large effective area ($>300\text{ cm}^2$) and a high spatial resolution (0.1 arc-sec) and imaging in far ultraviolet spectral range (110-190 nm) from a space telescope. The main part of the science payload is a Cassegrain-type telescope with a 60 cm-diameter primary mirror. As a current design, three main UV instruments are installed on the focal plane of the telescope: a mid-dispersion spectrograph, a high-dispersion spectrograph, and a slit imager. The mid-dispersion spectrograph contains a movable slit with different slit width, a holographic toroidal grating with 2110 lines/mm, and an MCP detector coupled with CMOS imaging sensors. Spectral resolution of 0.02 nm and field-of-view of 100 arc-sec will be achieved. The high-dispersion spectrograph consists of a slit, a toroidal mirror, an echelle grating, a cross disperser, and a detector. Highest spectral resolution of 3 pm will be achieved at the target wavelength (130.5 nm). The UV slit imager consists of imaging optics, several bandpass filters with a rotation wheel, and a same type of UV detector as the one installed in the spectrometer. In order to achieve a high spatial resolution of 0.1 arc-sec, we will install a target monitoring camera at 0th order position inside the spectrometer and slit imager for both attitude control and image accumulation process. We also plan to install a fine guidance sensor to monitor guidance stars. In addition, new key technologies such as funnel-type MCPs and CMOS-coupled readout system and highly reflective UV coatings will be applied to satisfy requirements of LAPYUTA. We have demonstrated performance of prototype models and samples of such key components. These key technologies can be applied to the future international flagship missions such as Habitable Worlds Observatory. Here we present the LAPYUTA concept design, the overview of the mission instruments, and the updated status of key technology developments.